

Tropospheric Emissions:  
Monitoring of Pollution



# Observing Nightlights from Space with TEMPO

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Brian D. Baker<sup>3</sup> and  
Kelly Chance<sup>2</sup>

**September 27, 2016**

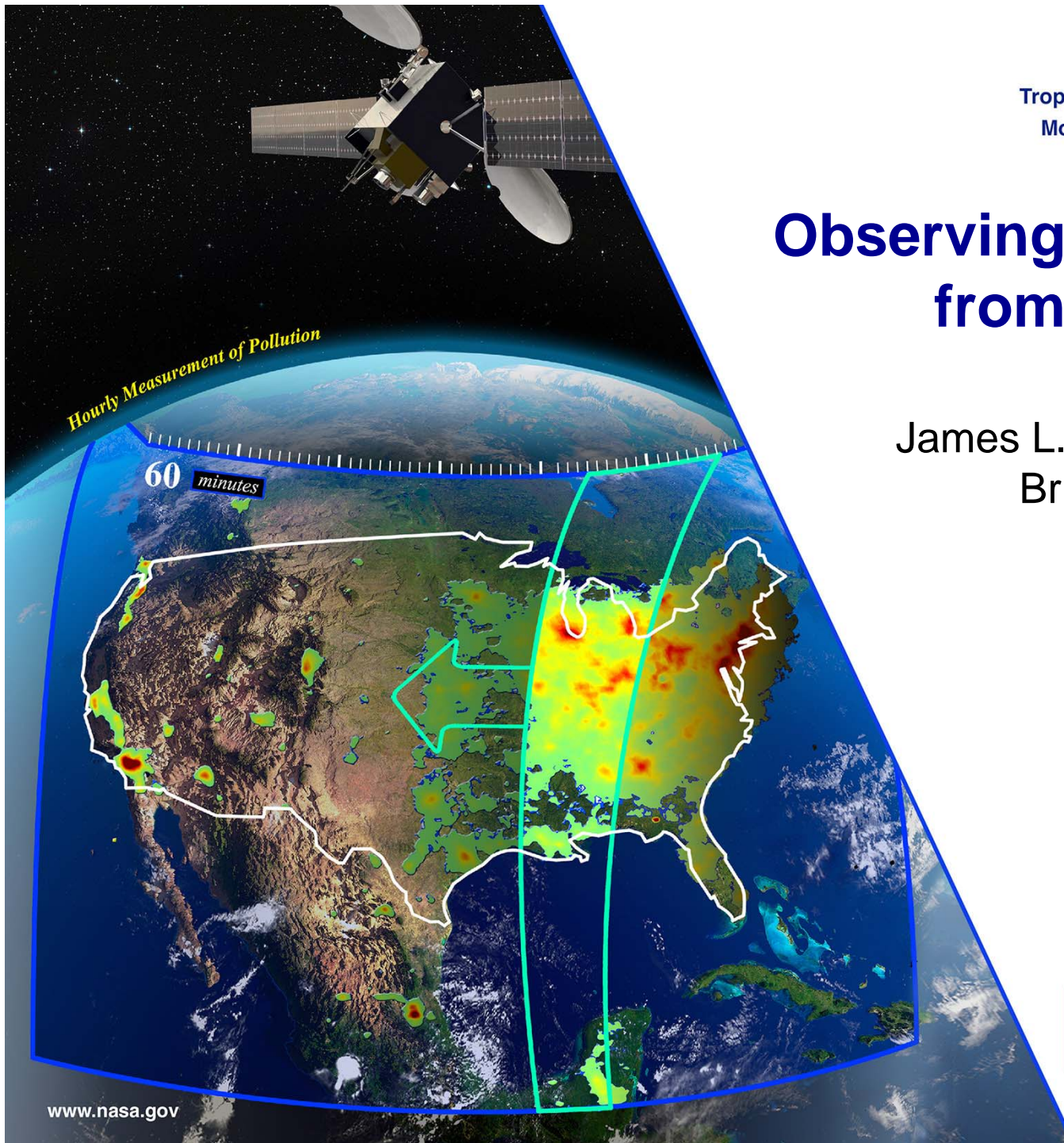
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Boulder, CO, USA*



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## ➤ **Tropospheric Emissions: Monitoring of Pollution (TEMPO)**

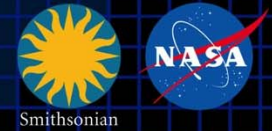
- NASA Earth Venture Instrument (EV-I)
  - Hosted on a geostationary commercial communications satellite over North America
  - Launch date TBD: 2018 – 2021
- Primary Mission: Air Quality / Atmospheric Chemistry
- Lead Institution: Smithsonian Astrophysical Observatory (PI: Kelly Chance, Deputy PI: Xiong Liu)
- NASA Implementing Center: NASA Langley Research Center
- Instrument Manufacturer: Ball Aerospace & Technologies Co.

## ➤ **Future International Constellation for Air Quality Observations**

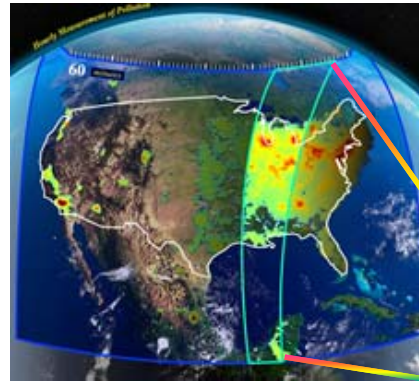
- Geostationary Orbit:
  - Geostationary Environment Monitoring Spectrometer (GEMS) – Korea, 2019
  - Sentinel-4 – Europe, 2022
- Low-Earth Orbit: Sentinel-5p (TROPOMI, 2017) and -5 (2021)



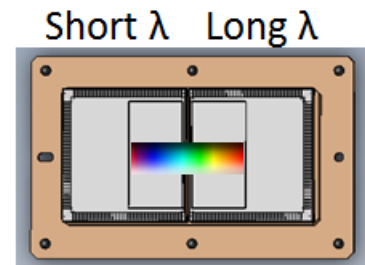
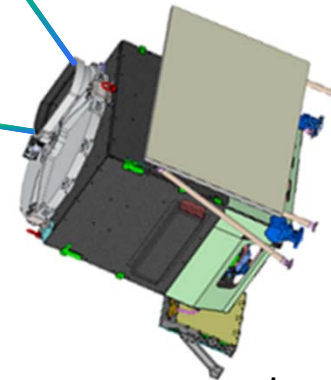
# TEMPO Operations: Step / Stare Imaging over Field of Regard



Parameter	Current Best Estimate
Frame Integration Time	118 ms
Image Frame Rate	7.92 Hz
Image Frame Time	2.65 s
Number of Coadds	21
Scan Mirror Step Size	114 $\mu$ rad
Number of Scan Mirror Steps	1283
Coverage Time	59.1 min

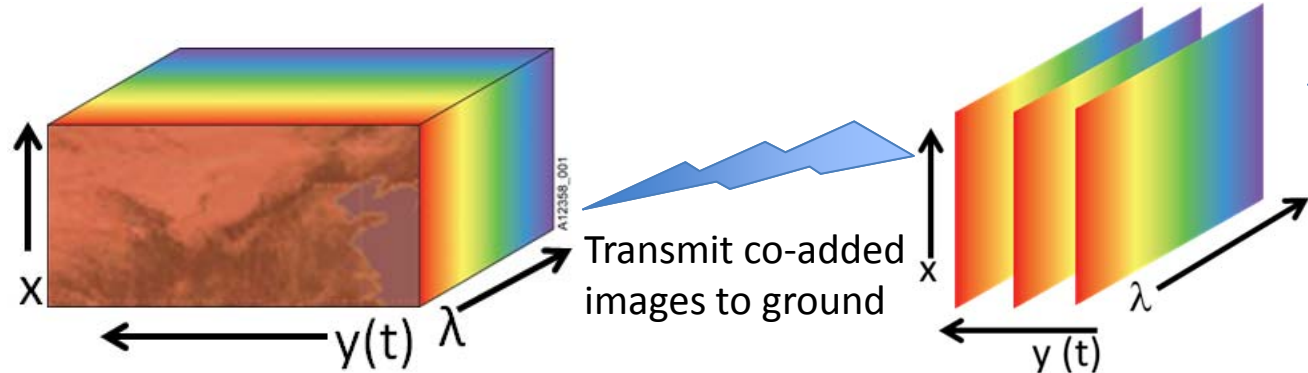


TEMPO step / stares over "Greater North America" in 1283 steps from East to West over 59.1 minutes



Ground processing spatially bins and geo-locates image

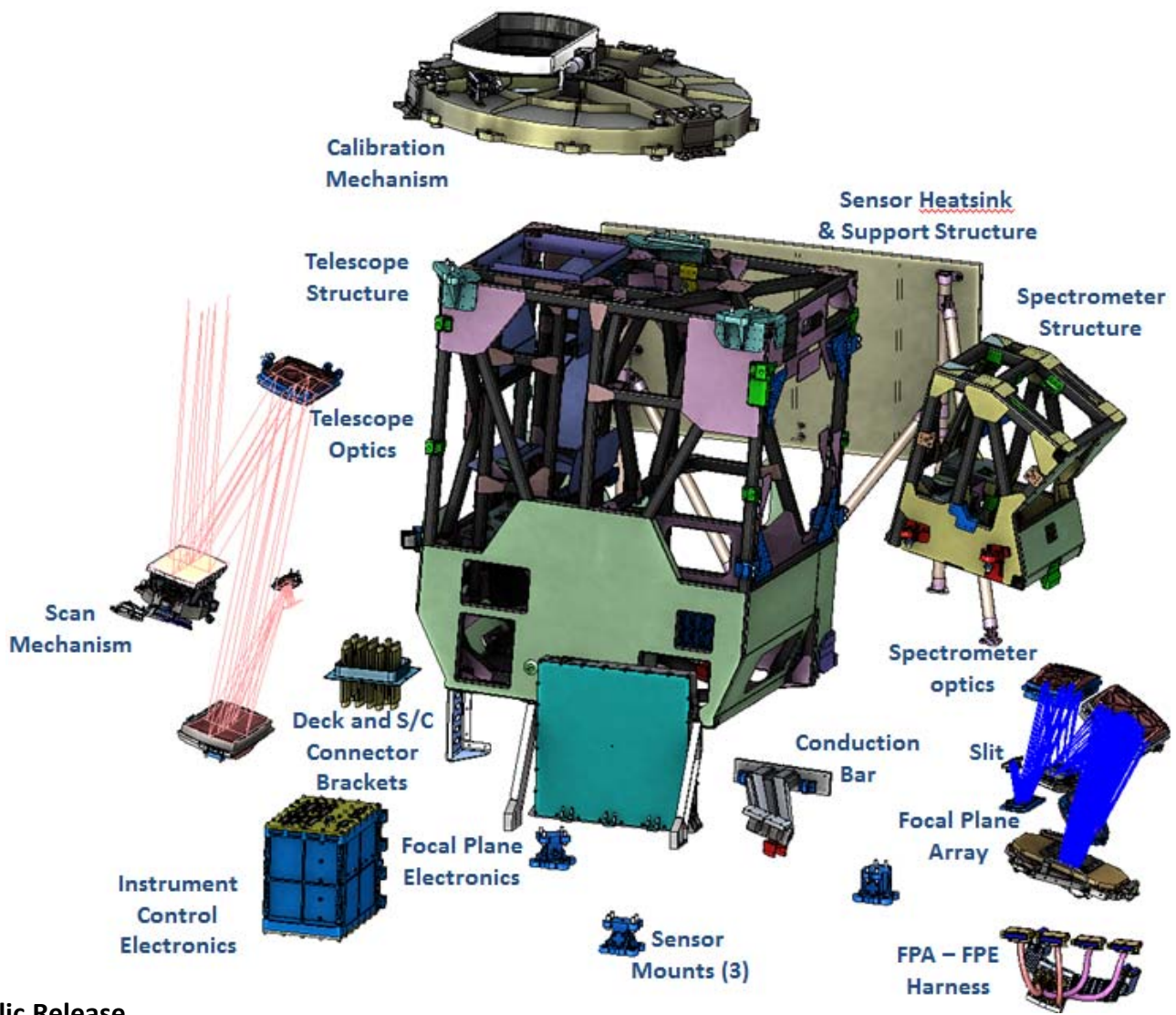
Images from each scan mirror position are co-added on board



Approved for Public Release



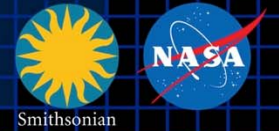
# TEMPO Instrument: Expanded View



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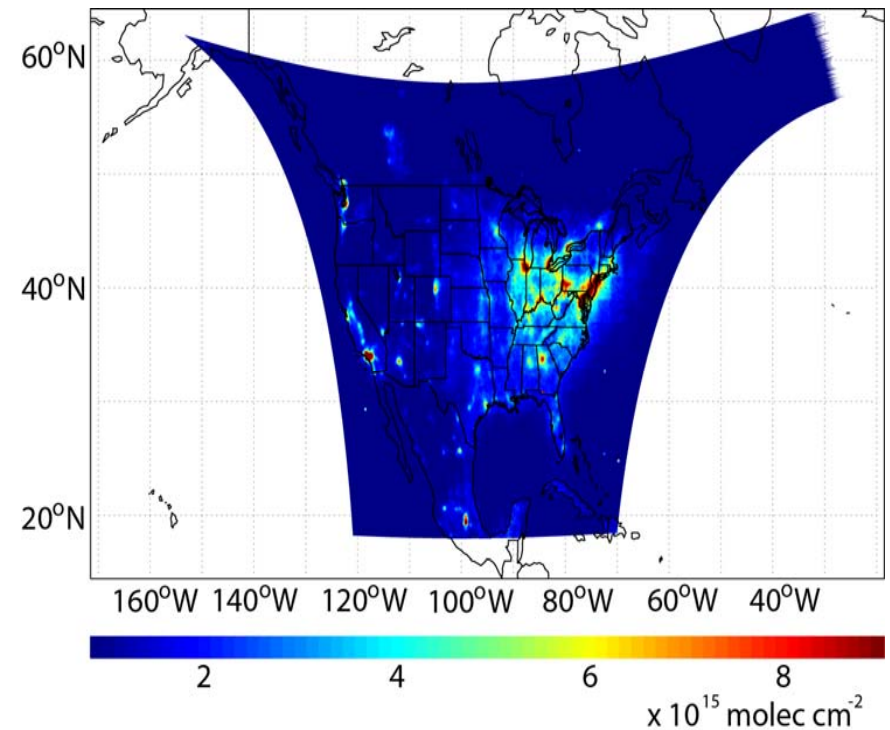


# Daytime Mission



- Cover Greater North America from a geostationary orbital station 80°W to 115°W
- Retrieve concentrations of trace gases, including O<sub>3</sub>, NO<sub>2</sub>, SO<sub>2</sub>, H<sub>2</sub>CO, C<sub>2</sub>H<sub>2</sub>O<sub>2</sub>, BrO, IO, and H<sub>2</sub>O from the spectra of reflected sunlight, plus other objectives (aerosols, clouds, chlorophyll fluorescence)
- Revisits < 1 hour during daylight to follow the creation and dispersal of pollutants
- Ground Footprint: 2.1km x 4.4km at Field of Regard center from GEO at 100°W
- Dwell per pixel: 2.478 s

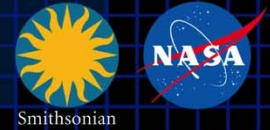
P. Zogman, et. al., "Tropospheric Emissions: Monitoring of Pollution (TEMPO)", *Journal of Quantitative Spectroscopy and Radiative Transfer*, in press.



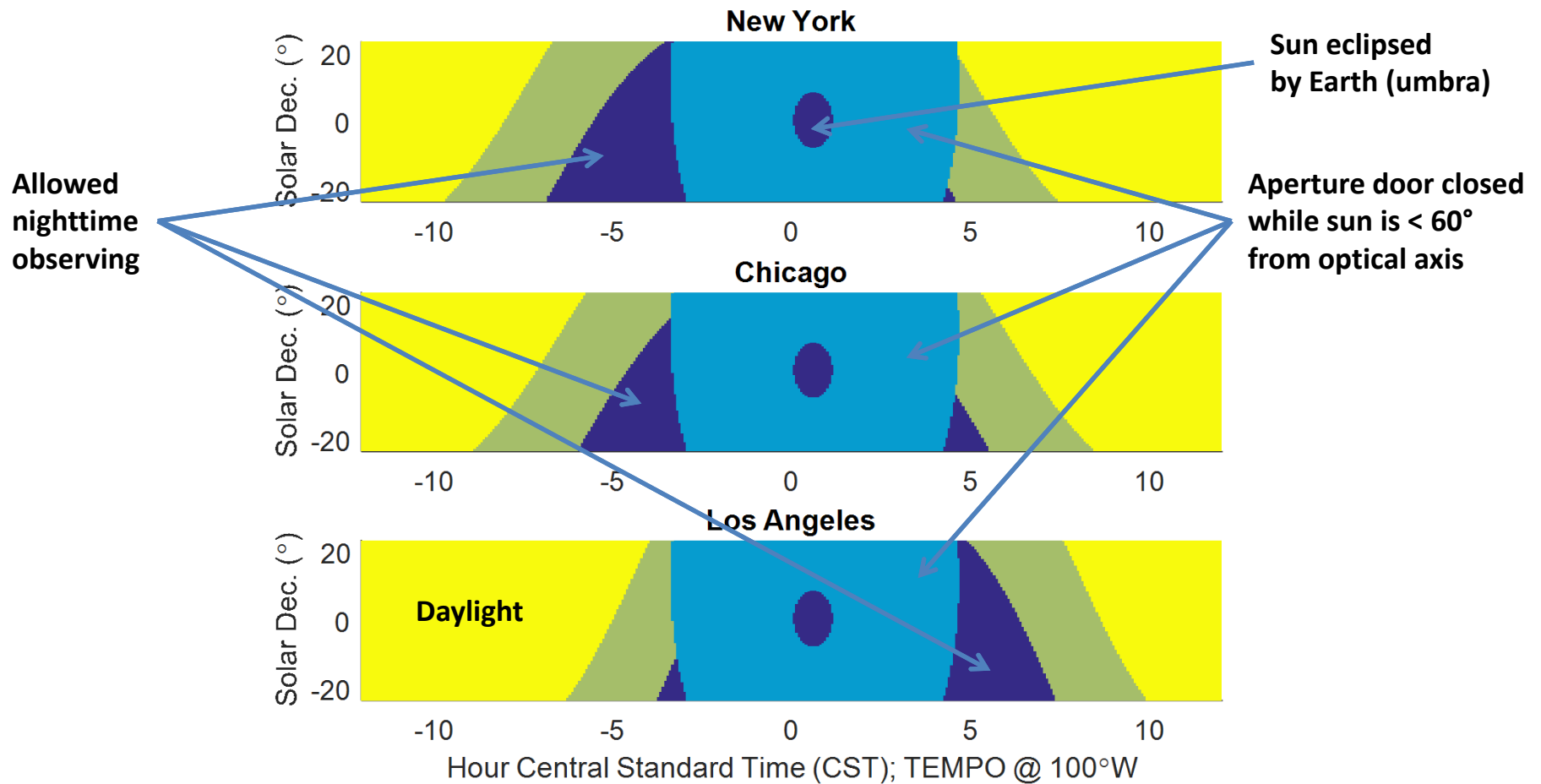
**NO<sub>2</sub> column densities over the TEMPO Field of Regard (derived from OMI)**



# Nighttime Observing

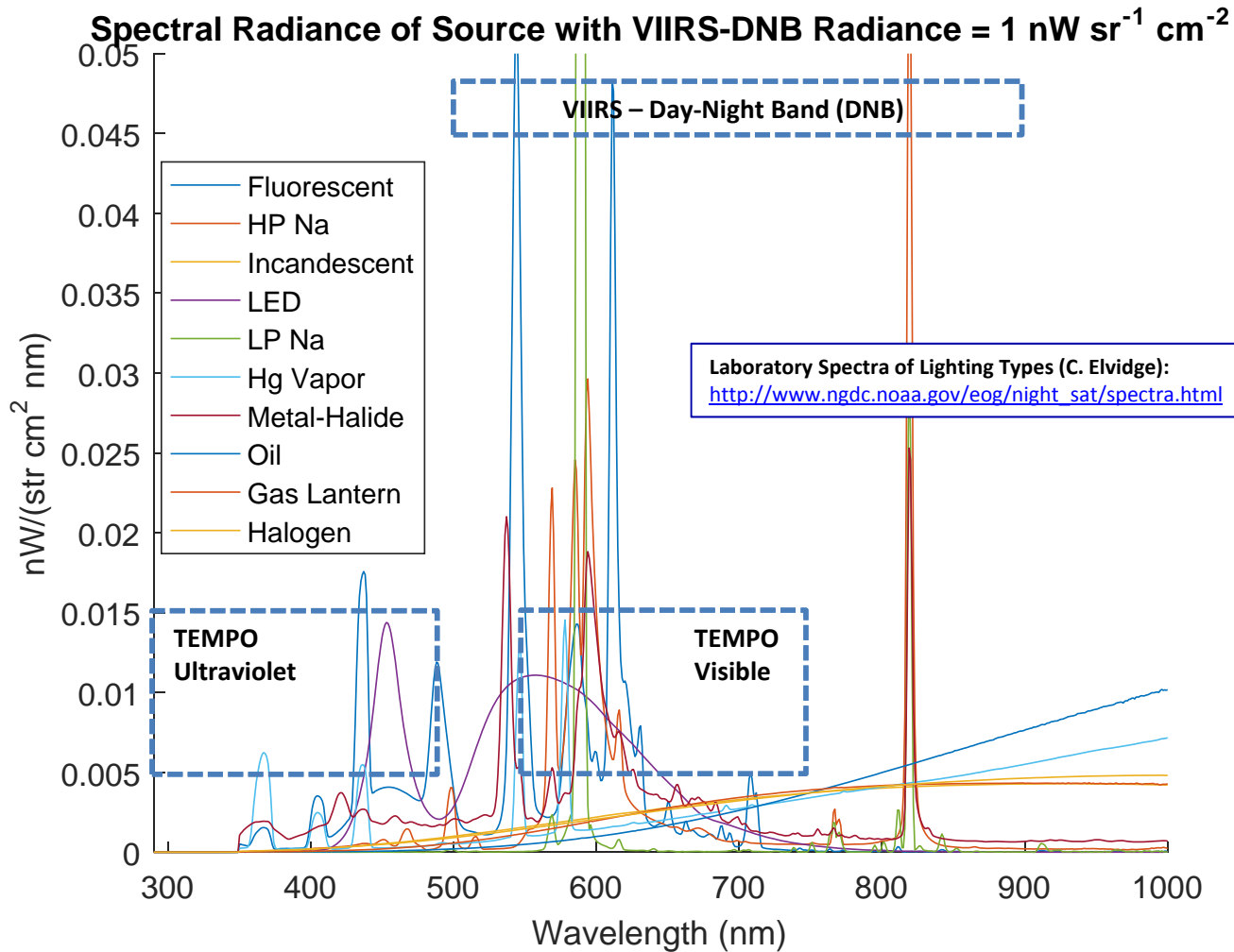
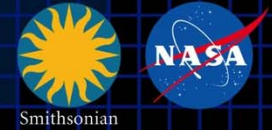


- Sun safety constraint limits nighttime observing opportunities
- Best times for TEMPO nighttime observations are during winter



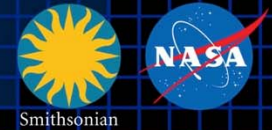


# Spectroscopic Signatures





# Nightlight Retrievals

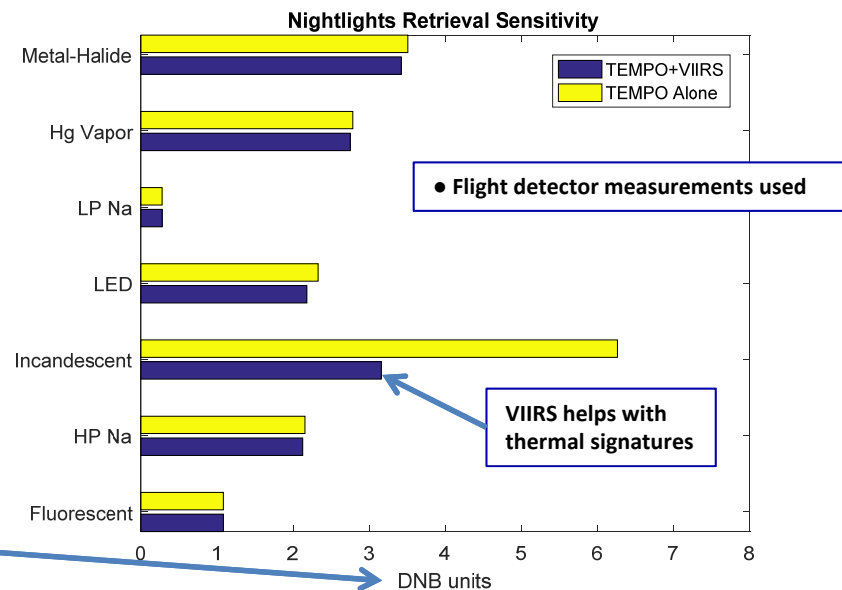


## ➤ TEMPO nighttime observing Operations Concept

- Plan for clear skies over areas of interest & sun constraint
- Increase dwell time to ~10s per pixel
- Calibrate dark current with aperture shut 16x 10 dwells
- Open aperture and collect over designated area
- Recalibrate dark current with aperture shut ~160s
- Leave aperture shut while waiting to resume daylight operations

- Nightlights spectral fitting “retrieves” nightlight radiances from each lighting type in our library and accounts for moon
- Poisson noise in TEMPO dark current is limiting factor

Uncertainty in retrieved radiances are in units of an equivalent response in the VIIRS-DNB. One “DNB” unit =  $1 \text{ nW sr}^{-1} \text{ cm}^{-2}$ .



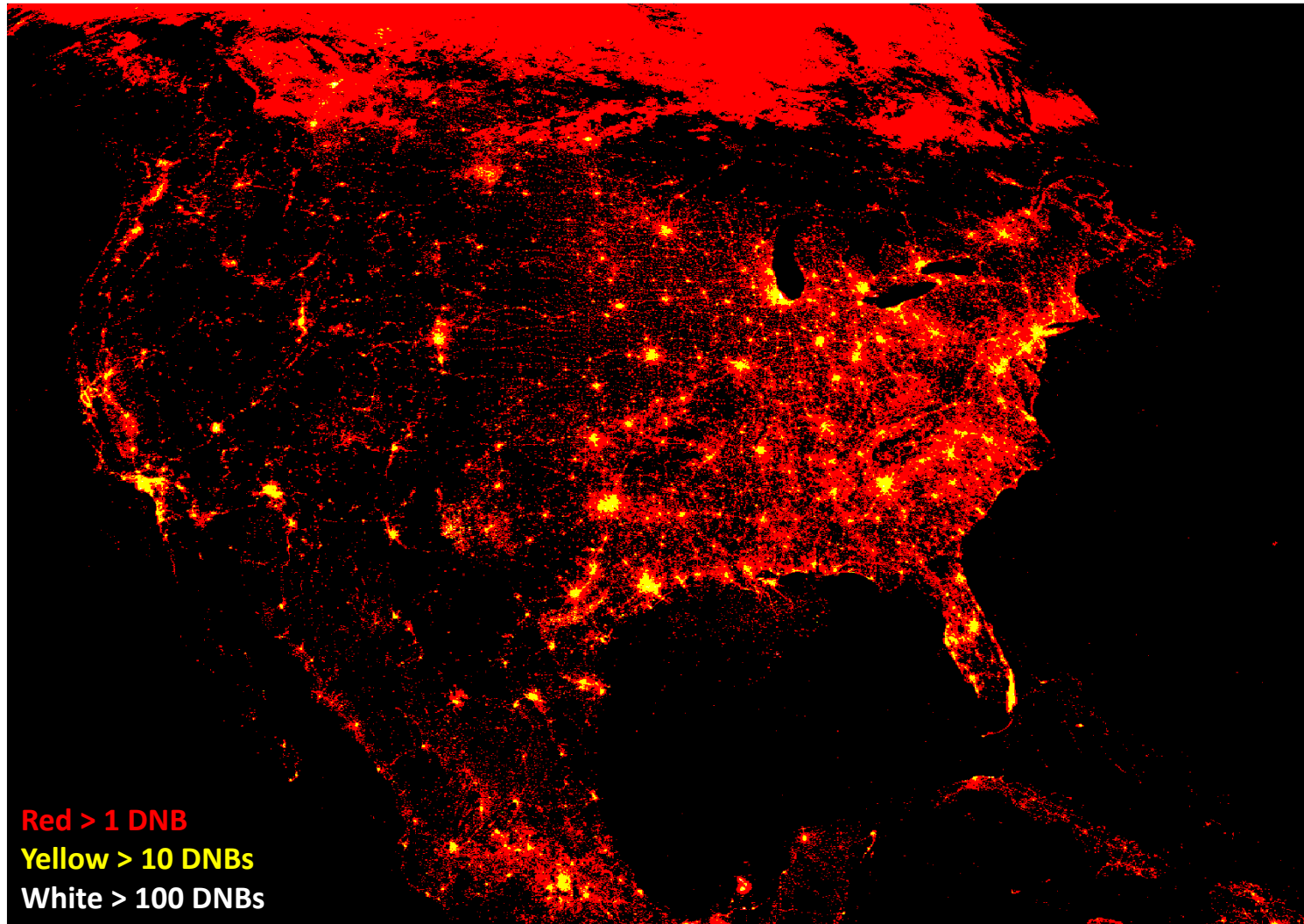




# What TEMPO Might See

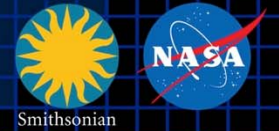


December 2015 Clear-Sky Mean VIIRS-DNB Radiances (NOAA/NGDC)  
Remapped to the TEMPO Field of Regard and Resolution





# Conclusions



- Fine spectral resolution of the TEMPO spectrometer enables discrimination of different lighting types that is simply not possible using the VIIRS-DNB
- TEMPO nightlights retrievals will be sufficiently sensitive to characterize outdoor lighting types over North America
  - Enhanced pixel dwell times for low light
  - Sensitivity is limited by Poisson noise in dark current
  - Most sensitive to sources with highly structured spectra
  - Best observing is during winter at Beginning of Life
- We encourage our Korean and European colleagues to look at the capabilities of GEMS and Sentinel-4 to do similar exciting new science with their instruments

