# A proposed method for estimating regional and global changes in energy consumption for outdoor lighting

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$$\Delta E = \frac{s\mu_{s1}L_{s1}K_{s1} + (1-s)\mu_{p1}L_{p1}K_{p1}}{s\mu_{s2}L_{s2}K_{s2} + (1-s)\mu_{p2}L_{p2}K_{p2}}$$

## Thanks to ...









#### Remote Sensing

of

## Energy Consumption for Outdoor Lighting

#### Light spectrum carries information

Visible Infrared



NASA/JPL-Caltech/R. Hurt (SSC)

#### Broadband radiance/reflectance



## Remote sensing

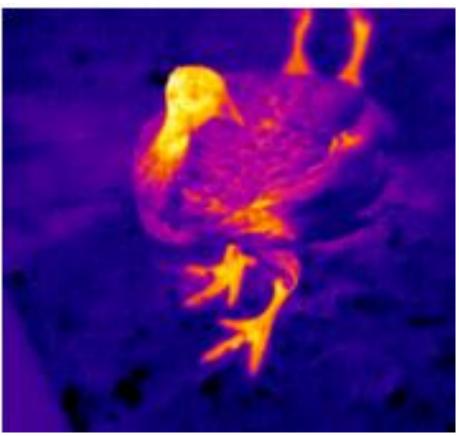


### Remote sensing

"Remote sensing is the acquisition of information about an object or phenomenon without making physical contact with the object..."

-Wikipedia





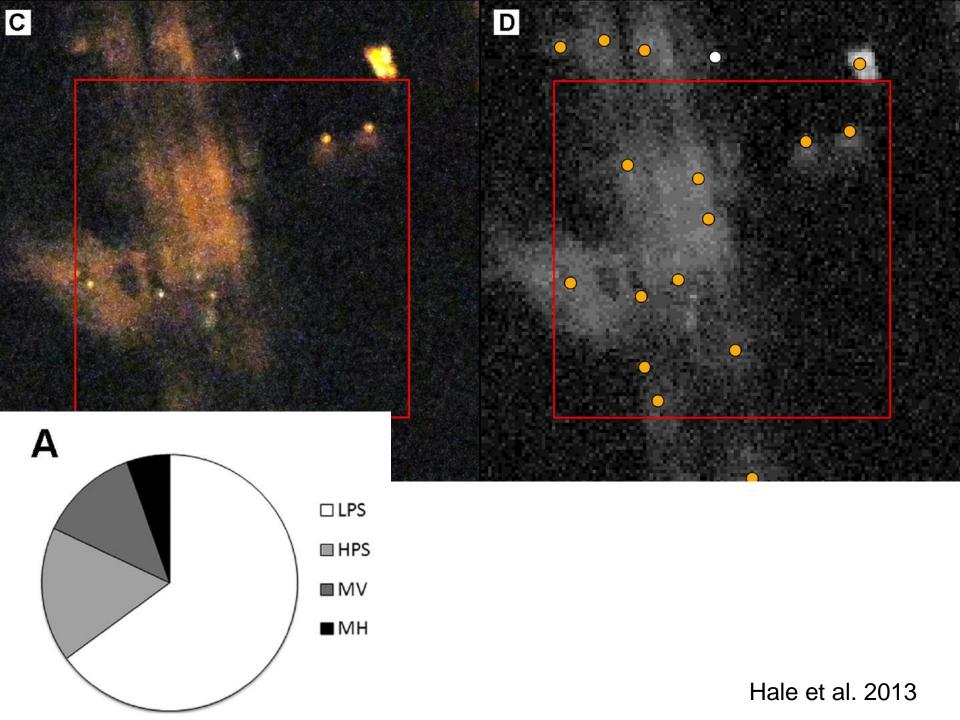
John Snell, Fluke blog www.thesnellgroup.com

### Remote sensing

"In current usage, the term 'remote sensing' generally refers to the use of satellite- or aircraft-based sensor technologies to detect and classify objects on Earth"

-also Wikipedia

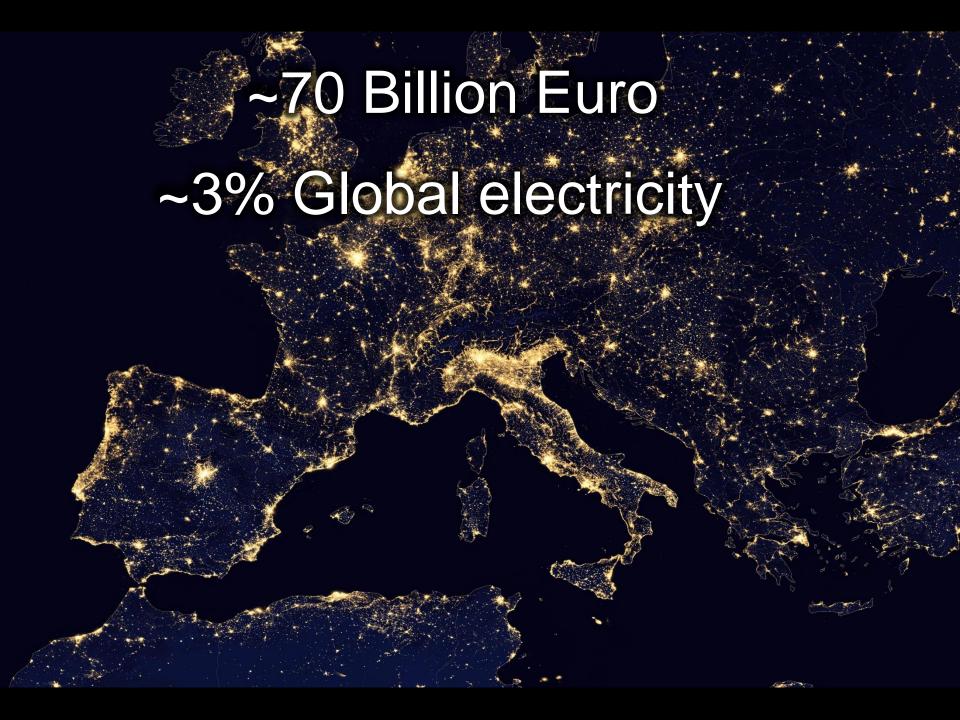




#### Remote Sensing

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## Energy Consumption for Outdoor Lighting

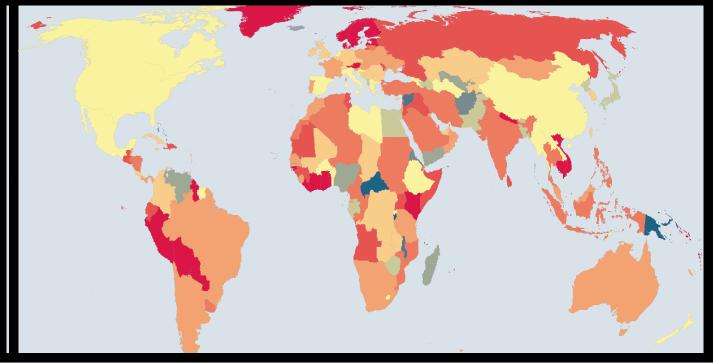


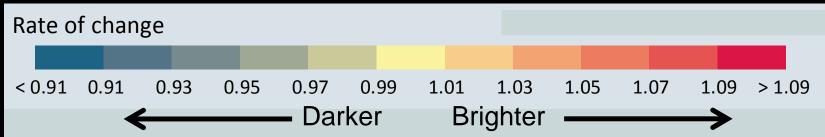
## Do LEDs save energy, or just enable brighter lights?

White LED

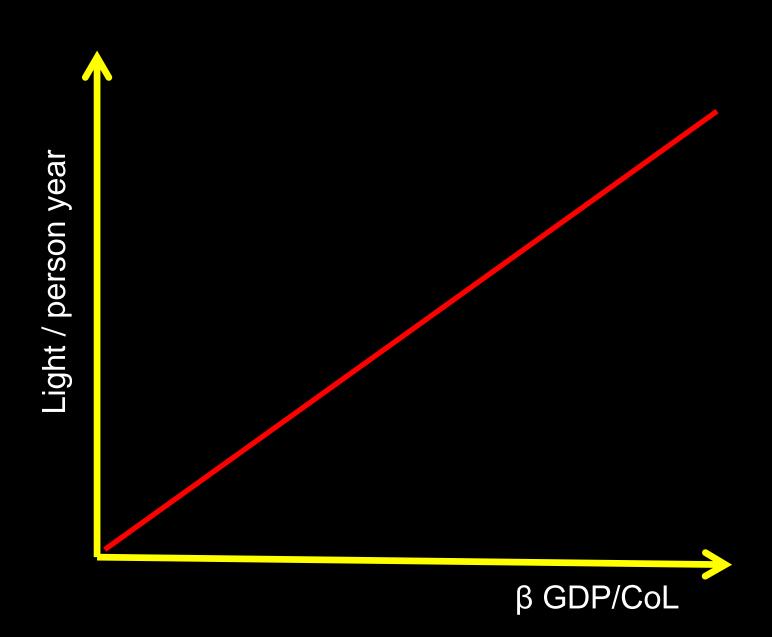
Sodium vapor

## Annual change in radiance (brightness), 2012-2016



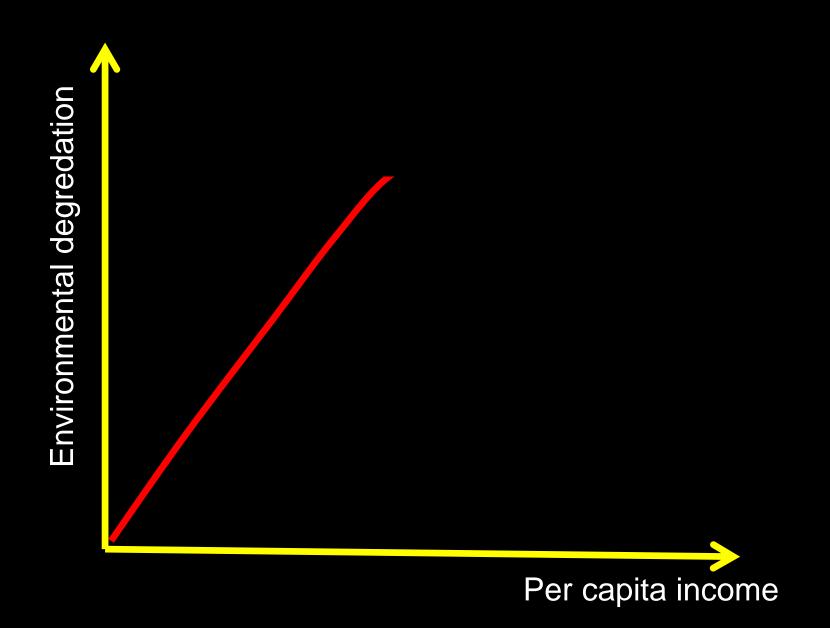


## Demand for lighting

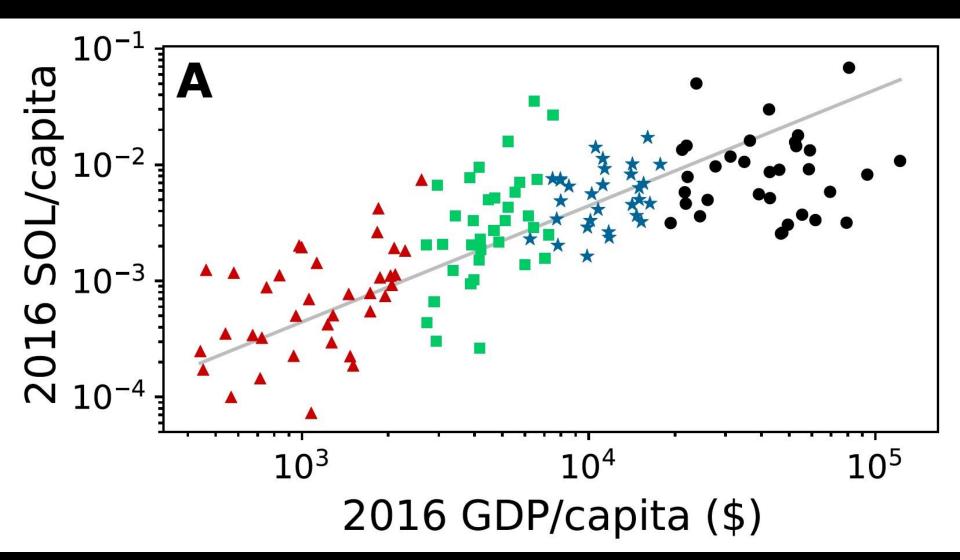


## Demand for lighting

### Environmental Kuznets curve



## Light/capita vs GDP/capita



#### Remote Sensing

of

## Energy Consumption for Outdoor Lighting

### Simple model

- All light is public outdoor area lighting
- Only one lamp type

## Berlin (VIIRS DNB)

- R radiance
- L luminance
- μL lumens
- K W/lumen
- t time

 $E = \mu LKt$ 



## Private lighting



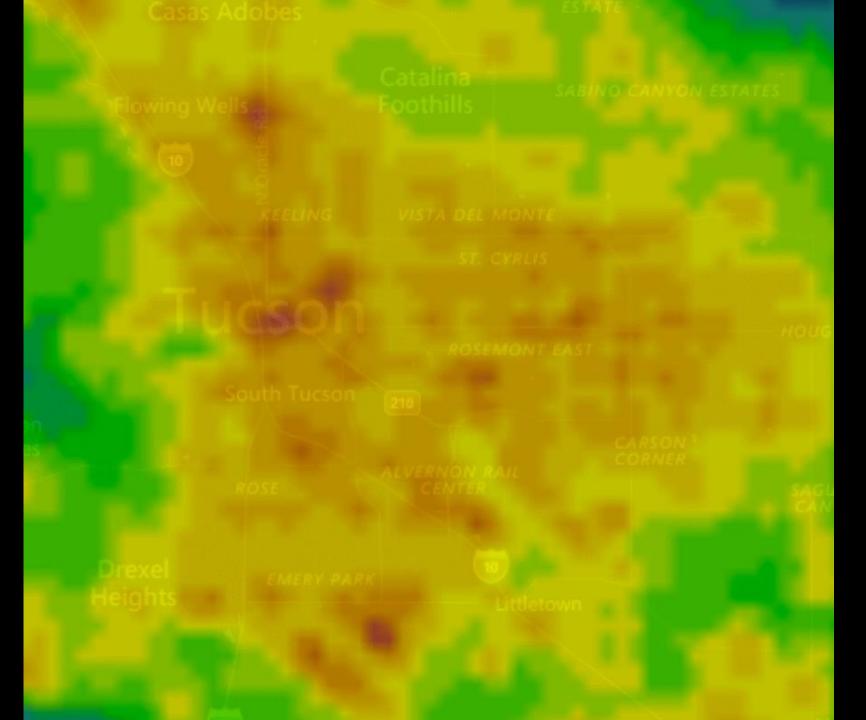
## City lighting change experiment



Full power

50% Dimmed

Photos of Cambridge, MA WSP | Parsons Brinckerhoff



R - radiance

L - luminance

μL - lumens

K = W/lumen

s – fraction streets

$$E = \mu LKt$$

$$E = s\mu_s L_s K_s t + (1-s)\mu_p L_p K_p t$$





#### You're welcome!

$$L_{s1} = \frac{L_1 - L_2}{1 - d} \quad , \quad L_{s2} = \frac{dL_1 - dL_2}{1 - d} \quad , \quad L_o = \frac{L_2 - dL_1}{1 - d} \quad , \quad s_1 = \frac{L_1 - L_2}{(1 - d)L_1}$$
 (3)

$$L_p(\hat{\Omega}) = \sum_n \alpha_n \sin \theta_n L_{p,n}(\hat{\Omega}, \theta_n)$$
 (5)

These predictions can be compared to observations  $L_o(\hat{\Omega})$  from an eVscope or all sky camera photograph:

$$\Delta(\hat{\Omega}) = L_o(\hat{\Omega}) - L_p(\hat{\Omega}) \tag{6}$$

A full eVscope acquisition or all-sky camera photo contains many observing directions  $\hat{\Omega}$ , so the overall agreement between the model and an observation can be assigned a test statistic, for example a likelihood  $\mathcal{L}_i$ 

$$\mathcal{L}_i = \prod_{\hat{\Omega}} p(\Delta(\hat{\Omega}))$$
 ,  $\mathcal{L} = \prod_i \mathcal{L}_i$  (7)